

Introduction

- **Nuclear Emissions** become **direct evidence** of LENR.
- For **Charged Particle** detection, (especially α)
possible environment is restricted.

↓
Gas permeation by **Mitsubishi Heavy Industries**

- **Disadvantages** of **CR-39** track detector
Time profile, Particle identification, E. calibration,
Background (cosmic rays, natural radioactivity)

Toward **extremely high precision** measurements

Charged Particle Detection

▪ Emissions are too weak in CMNS exp.

1. **Extremely low background**

(cosmic ray, natural RI)

(at least cosmic ray or not)

2. **Particle Identification**

3. High energy resolution

4. Time profile



Relation with exp. Condition

5. Large detection area (efficiency)

6. No shielding materials

7. Stable in **high temperature** hydrogen

8. Electrical noise

Ce doped Yttrium Aluminum Perovskite : **YAP(Ce)**

YAP(Ce) Characteristics

- Non-hygroscopic, Small temp. dependent (+2, -8 % ~200°C)
- **0.5mm thickness**, 20 mm ϕ , optical polish (Crytur, Czech)
- Max. Measuring E. (SRIM2008)
12.5MeV proton, 50MeV α

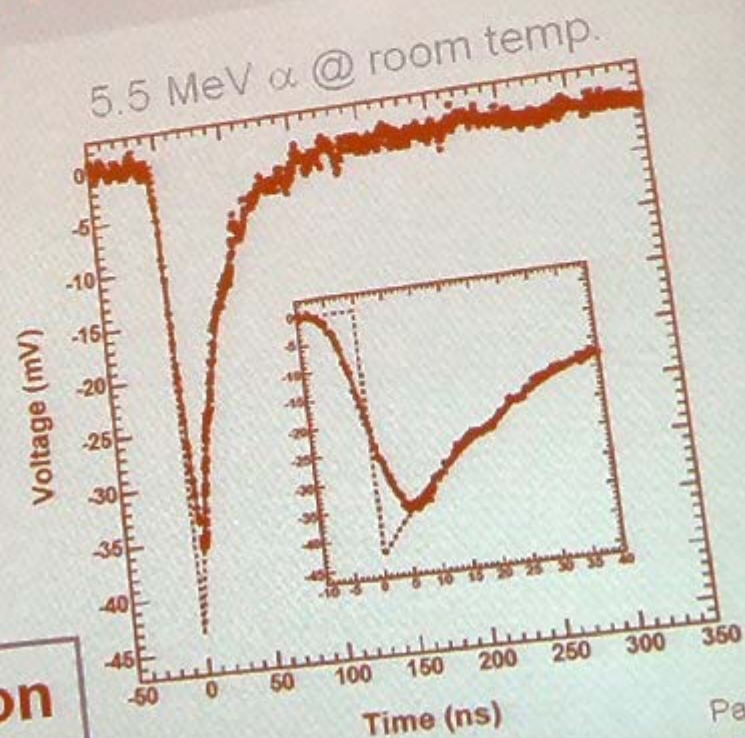
▪ **Resolution** : 22% in FWHM

▪ **Decay constant** : $\tau \sim 30$ ns

$$I = I_0 \exp\left(-\frac{t}{\tau}\right)$$

exhibits **particle identification**.

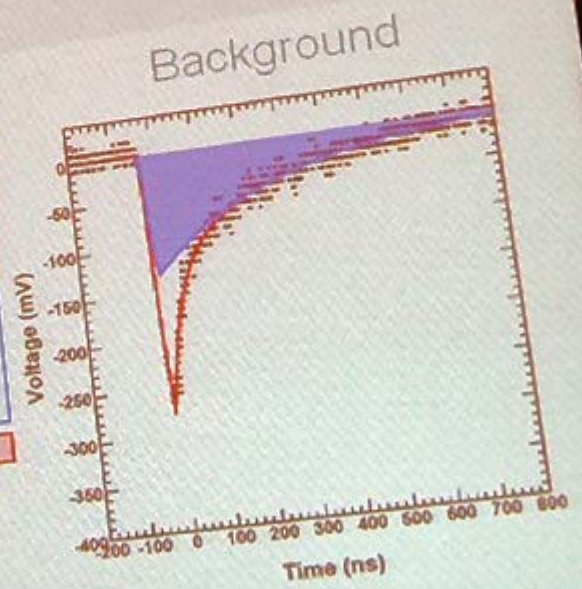
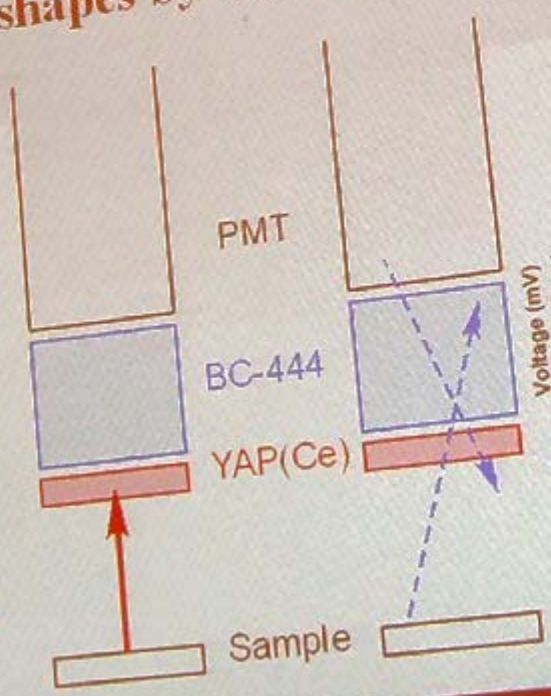
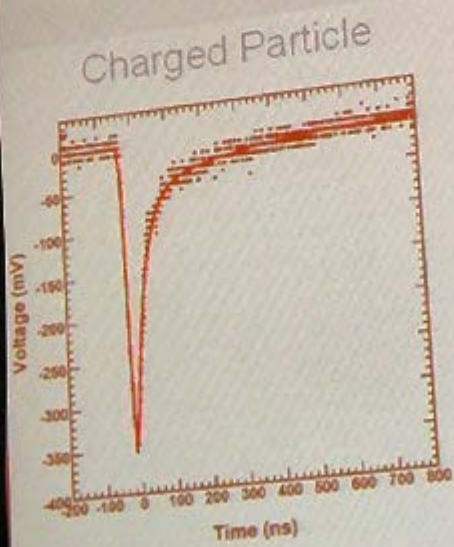
Pulse Shape Discrimination



Phoswich Detector

2 Scintillators / 1 PMT

- Front : **YAP(Ce)**, 20mm ϕ , 0.5mm, $\tau_{\text{YAP}} \sim 30\text{ns}$
- Back : **BC444 Plastic**, 20mm ϕ , 5mm, $\tau_{\text{BC}} \sim 300\text{ns}$
- We record **all pulse shapes** by digital storage oscilloscope.

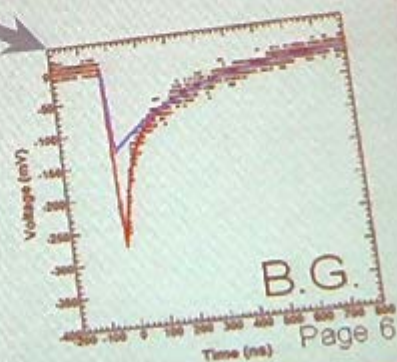
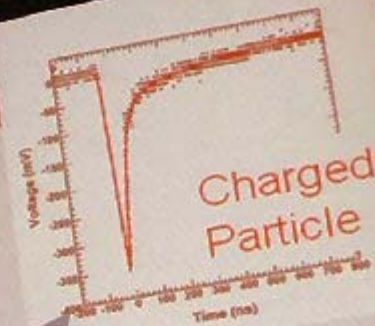
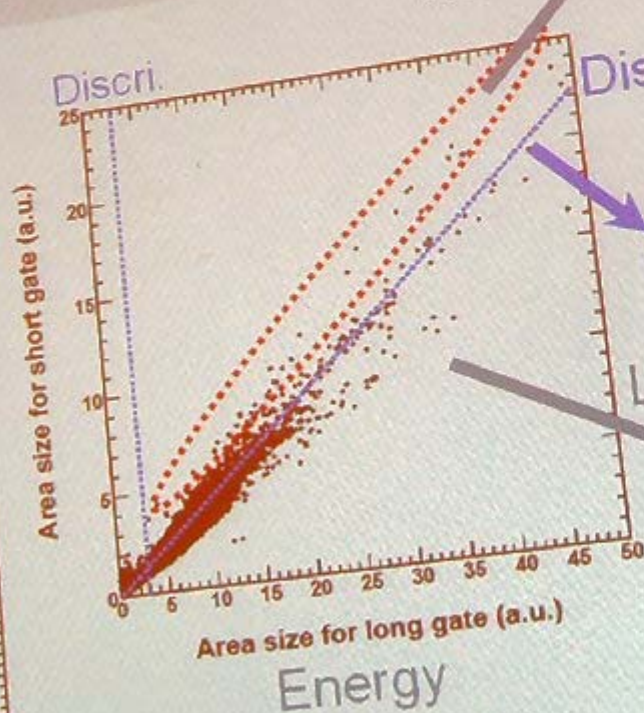
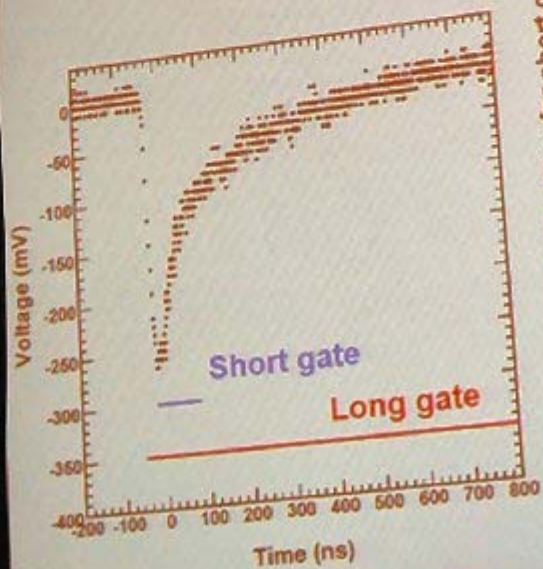


Pulse Shape Discrimination 1

2 integrations provide **Particle Identification.**

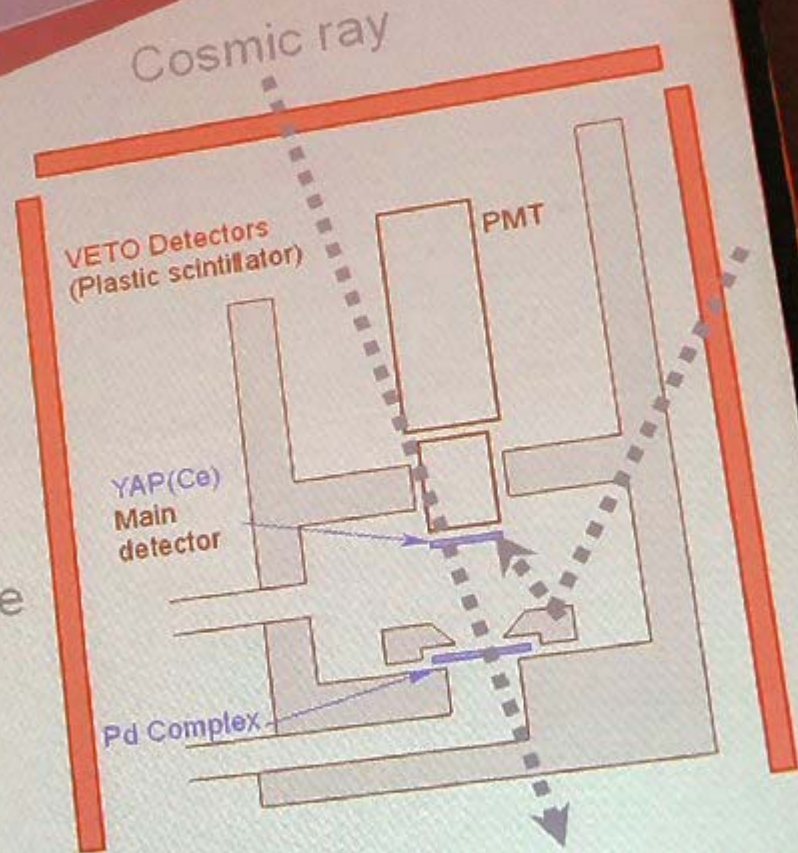
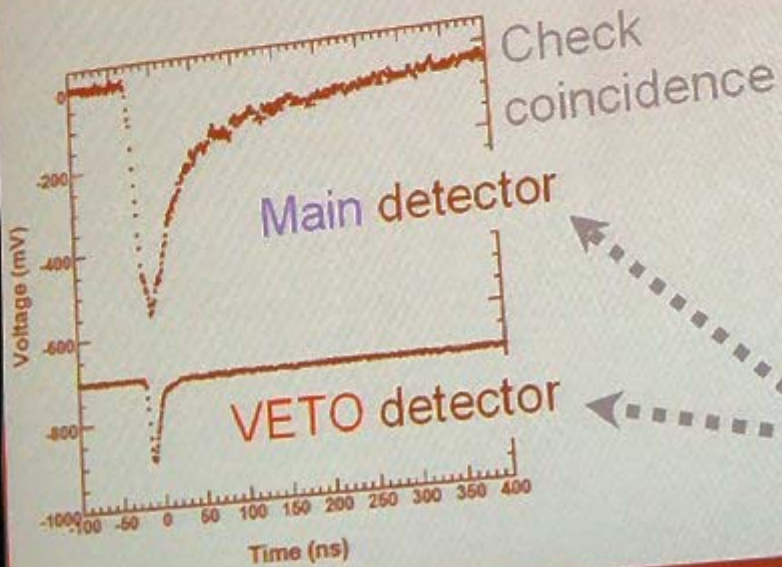
Short gate : 100ns

Long gate : 850ns



VETO Detectors 1

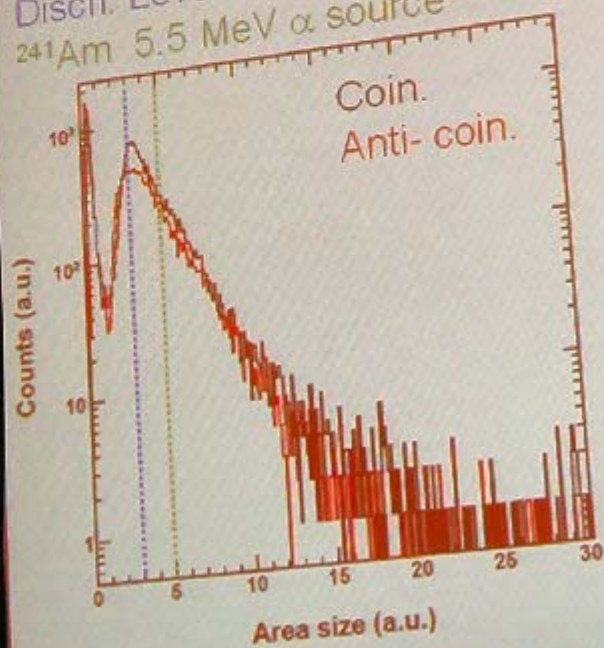
- Reduce cosmic ray events. (include secondary particle)
- **Natural radioactivity** cannot be suppressed.



Main & VETO detectors response at same time.

VETO Detectors 2

MHI sample,
D₂ permeation, 333.7 hours,
Discri. Level
²⁴¹Am 5.5 MeV α source



- **Coincident events**

100 % cosmic ray events
(ignore chance event)

- **Anti - coincident events**

Real events or cosmic rays
(not cover all area)

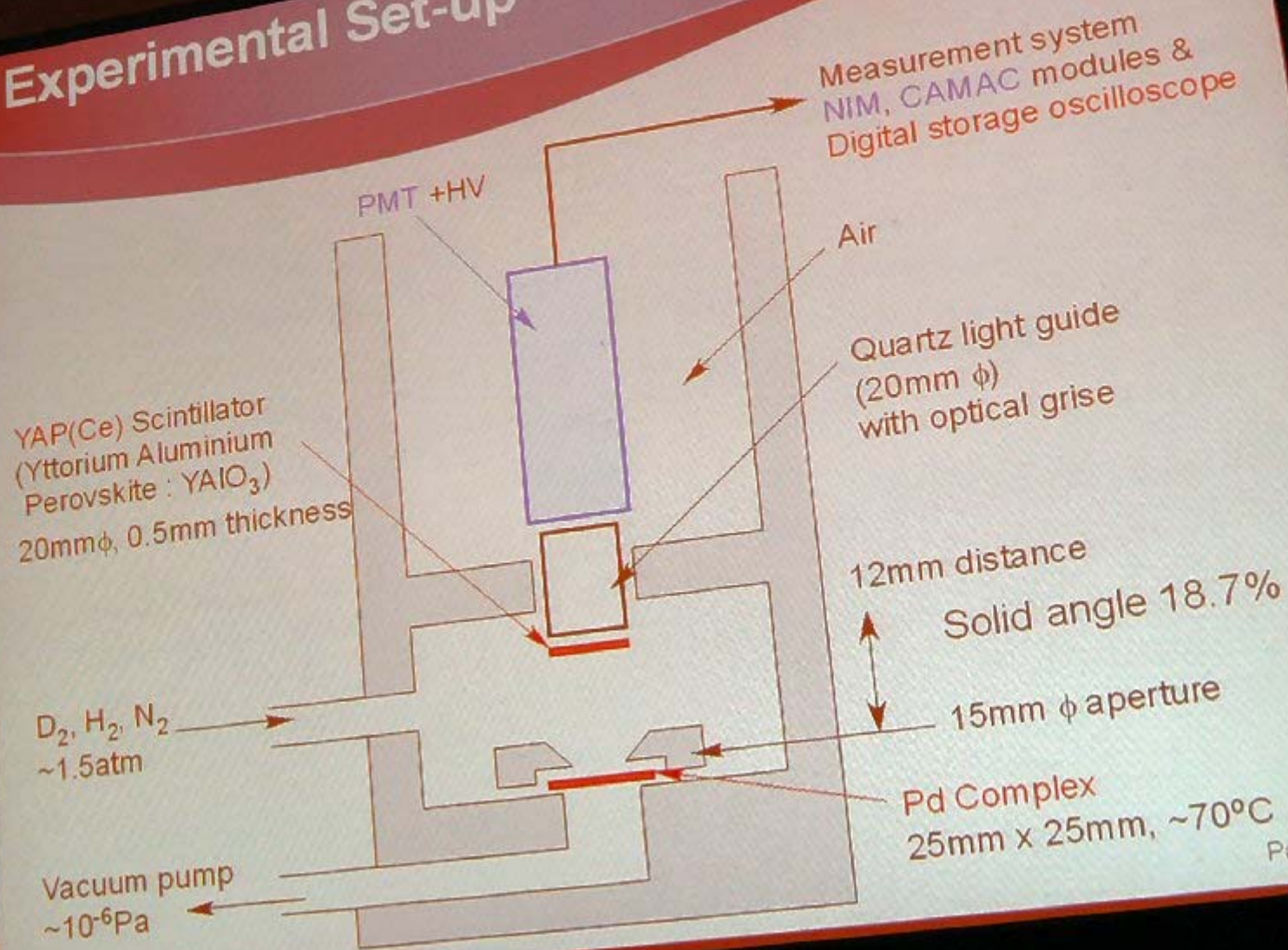
- **57 %** is rejected.

- **~75 %** could be predicted by additional detectors.

- Origin of **high energy** counts are unknown.

Elastic recoil ? $p(\mu, \mu)p, d(\mu, \mu)d$

Experimental Set-up



Picture

NIM, CAMAC
Modules &
Digital Storage
Oscilloscope

PMT & YAP(Ce)

D₂, H₂
gas

100 mm

Pd sample, 25 x 25 mm
Heater & T.C.

Vac., TMP

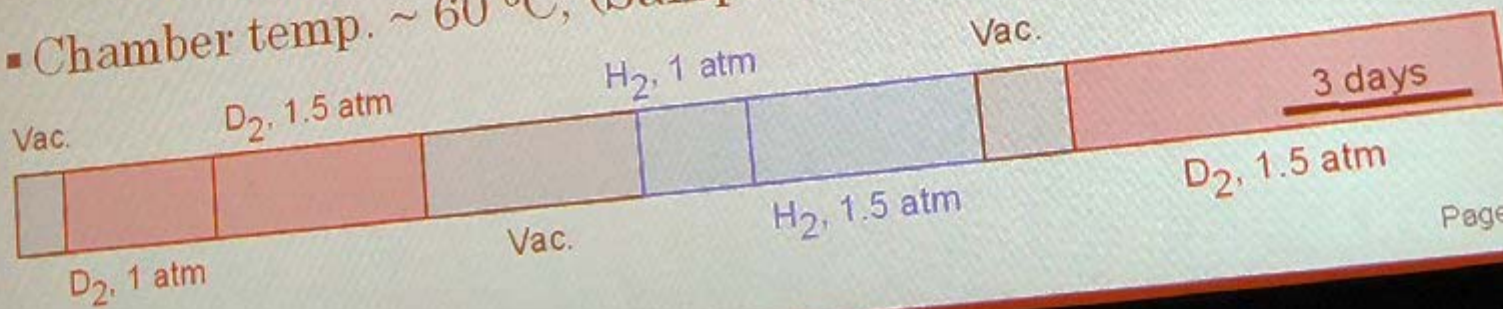
Chamber VETO Detector

Samples & Permeation

- Pd Complexes were supplied by **Iwate Univ. & MHI**
- Pd (0.1 mm) / CaO (2 nm) & Pd (18 nm) 5 layers / Pd (60 nm)
- Cs is electrodeposited. (Iwate sample)

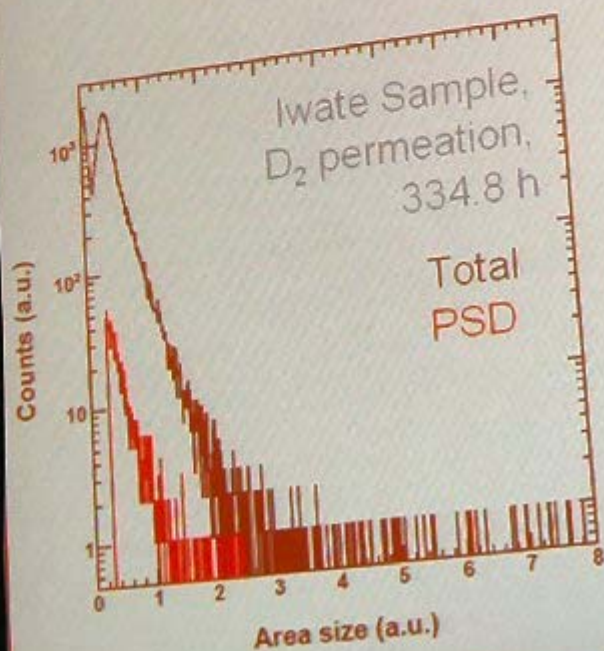
→ Yamada *et al.*, ICCF12

- Vac. B.G. & 1 ~ 2 weeks D₂ & H₂ permeation (1.0 ~ 1.5 atm)
- Flow rate ~ 0.5 ccm(Max.), typically 0.1 ccm
- MHI condition (over 1 ccm) has never been achieved.
- Chamber temp. ~ 60 °C, (Sample temp. is higher : ~70 °C)



Result 1 : Iwate Sample

- Reduction of B.G. (dark current, long tail event)
- 84684 counts in total $\xrightarrow{1.4\%}$ 1150 counts / 334.8 hours
without VETO

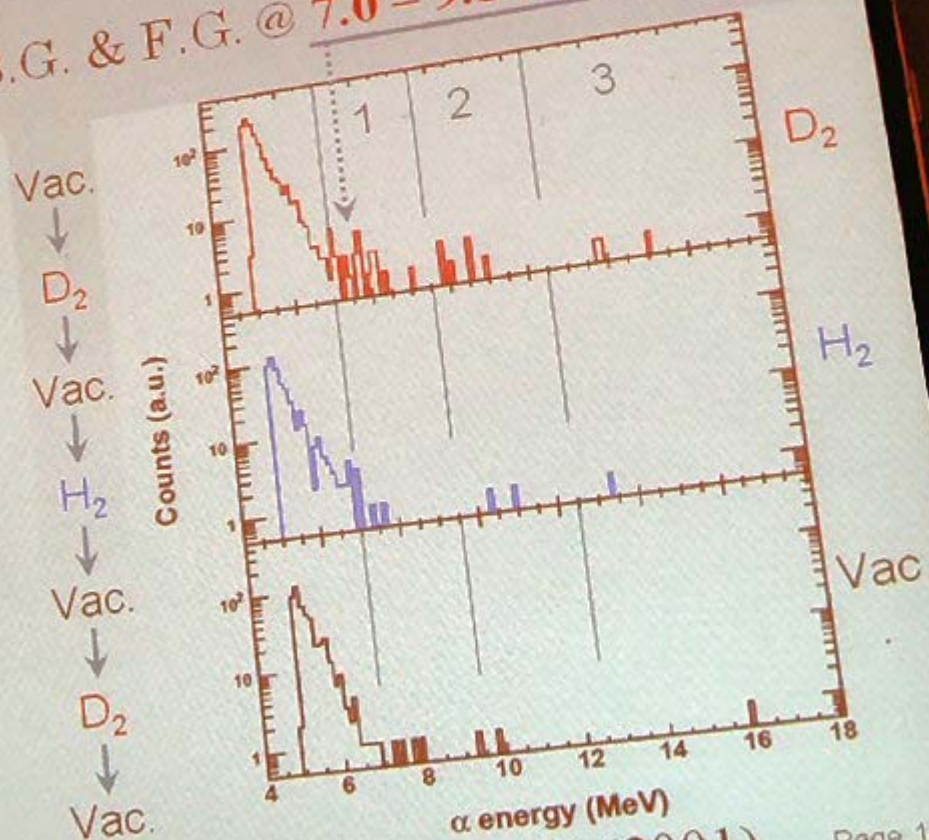
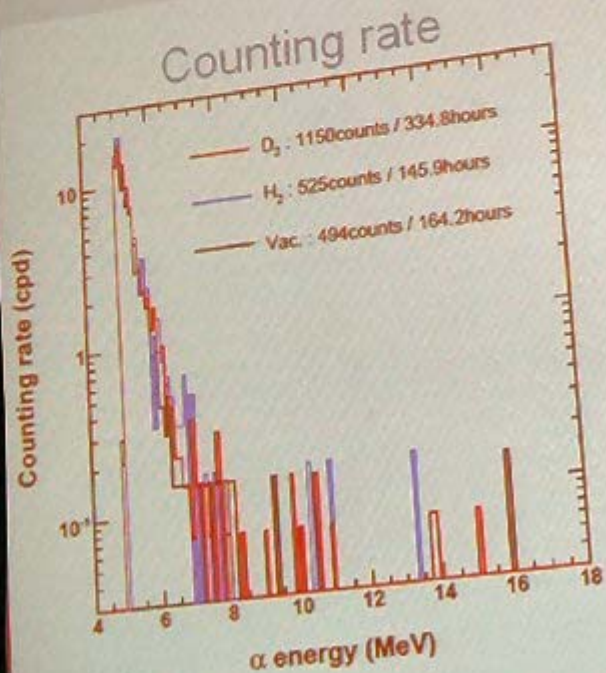


- Counting rate : ~ 3 counts / h
- Bad separation for low energy.
- Dark current & Cosmic rays are included in lower range.

▪ We can identify **3 counts / day** several MeV α particles.

Result 2 : Iwate Sample

▪ Small **difference**, between B.G. & F.G. @ **7.0 – 9.5 MeV**



Energy Calibration : *Nucl. Inst. Meth. A*, **469**, 70 (2001)

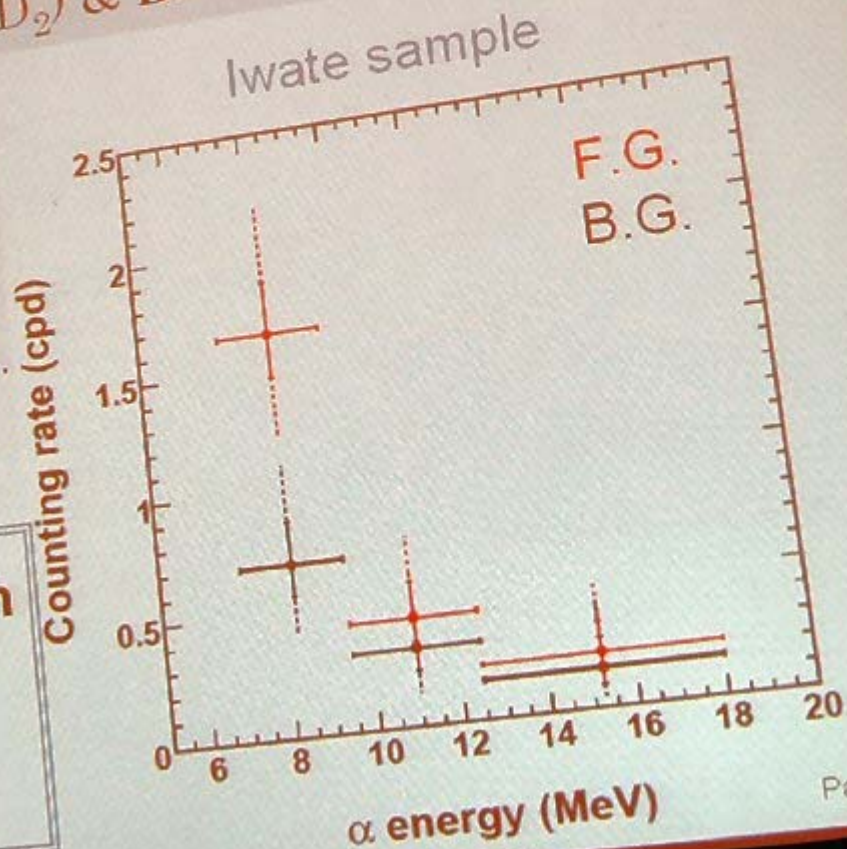
Final Result : Iwate Sample

- Counting rate [cpd] of **F.G.**(D₂) & **B.G.**(H₂ & Vac.)

- Vertical errors are
68.3 % (solid line)
90.0 % (dashed line)
confidence level.

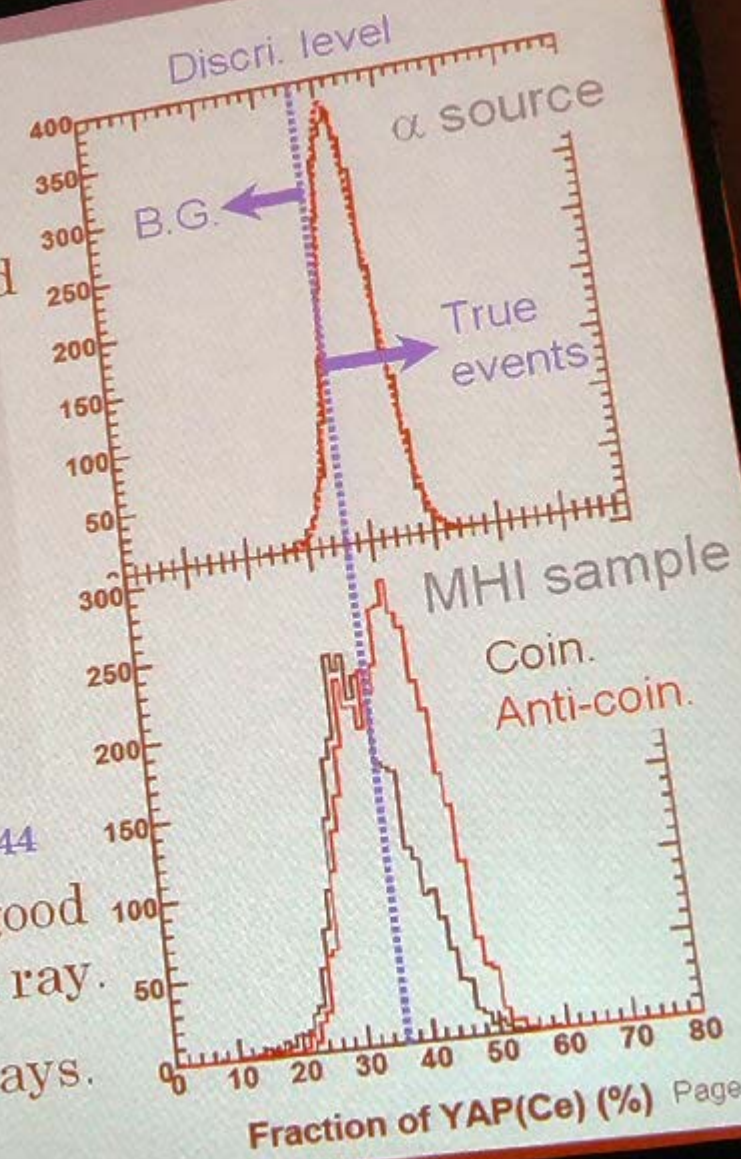
- At 7.0 – 9.5 MeV

**Charged Particle Emission
is Significant with
99.97 % reliability.**



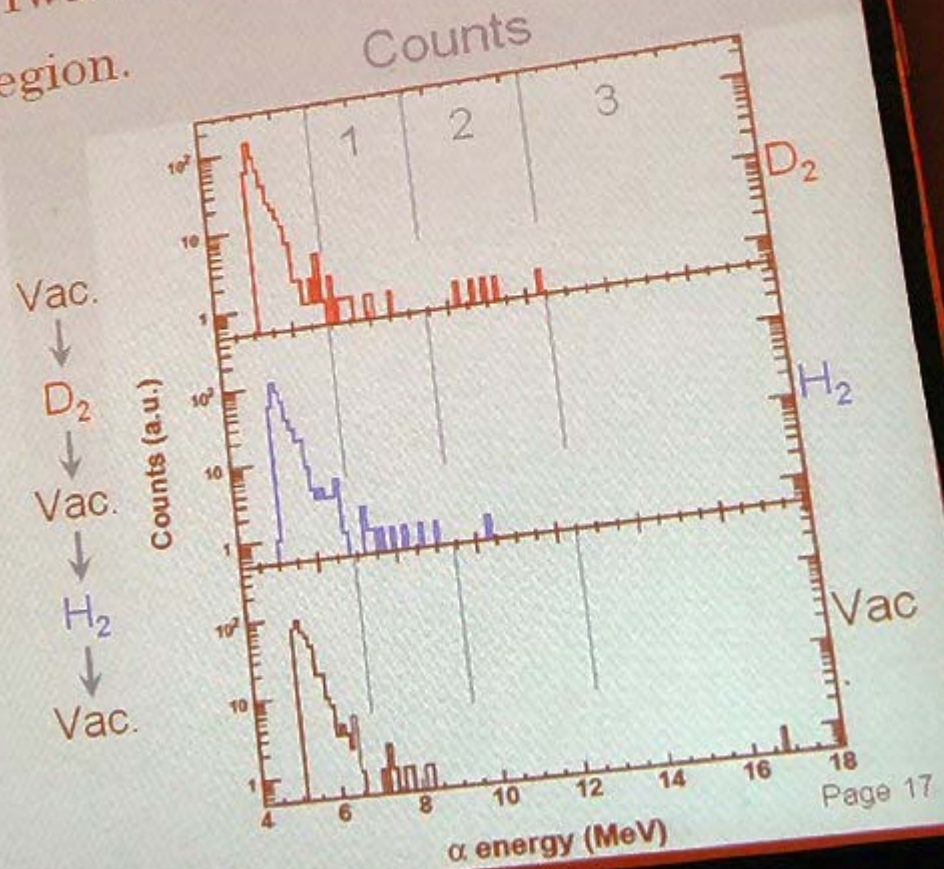
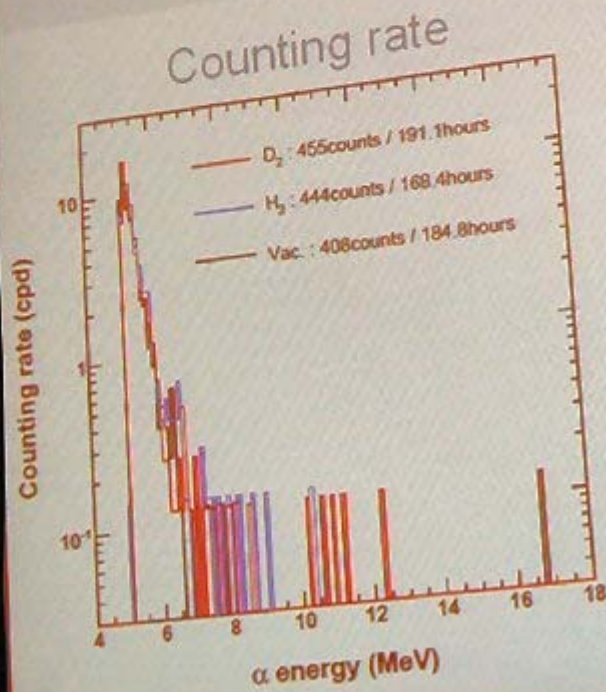
Result 3 : MHI Sample

- τ_{YAP} , τ_{BC444} & **Fraction** were derived by 2 component fitting.
- **Coin.** with VETO : All cosmic rays
- **Anti-coin.** : True events & **B.G.** (limited efficiency of VETO)
- **Discr.** : ^{241}Am source 1σ
- No significant difference for τ_{YAP} & τ_{BC444}
- **Fraction (Light yield ratio)** gives good separation, Cosmic / Not-cosmic ray.
- **PSD** can reduce almost cosmic rays.



Result 4 : MHI Sample

- Lower counts are similar to Iwate's.
- **Difference @ 9.5 - 12.5 MeV region.**



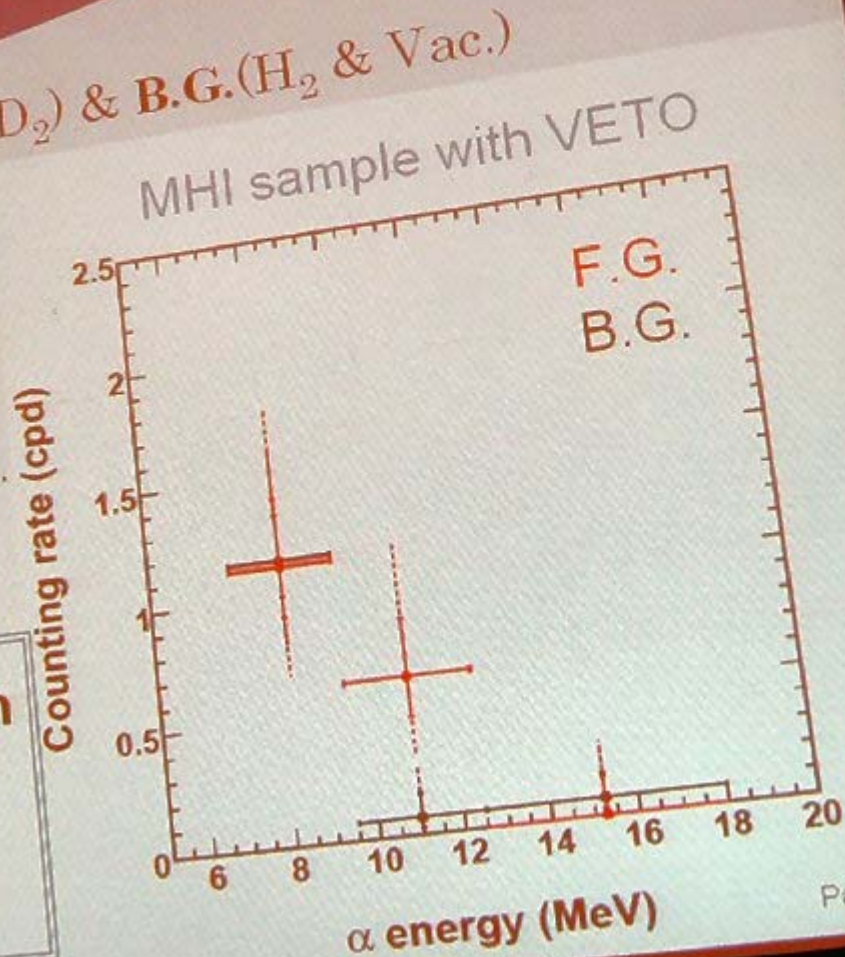
Final Result : MHI Sample

- Counting rate [cpd] of **F.G.**(D₂) & **B.G.**(H₂ & Vac.)

- Vertical errors are
68.3 % (solid line)
90.0 % (dashed line)
confidence level.

- At 9.5 – 12.5 MeV

**Charged Particle Emission
is Significant with
99.63 % reliability.**



Charged Particle Detections

High energy counts during D₂ gas permeation

Phoswich & PSD

▪ Iwate (Cs)

99.97 % reliability @ 7.0 - 9.5 MeV

Reaction rate : $R \sim 6.8 \times 10^{-3}$ [f/s/cc], $\sigma \sim 10^{-29}$ [barn]

▪ Iwate (no Cs)

No significant event

Phoswich & PSD & VETO

▪ MHI (Cs)

99.63 % reliability @ 9.5 - 12.5 MeV

$R \sim 4.4 \times 10^{-3}$ [f/s/cc], $\sigma \sim 10^{-29}$ [barn]

▪ MHI (Cs)

No significant event

but too **Bad Permeation** rate (almost 0)

▪ Significant Detection : 2 / 4 samples

Conclusion : Detector System

We developed a new detector system with **PSD** analysis, **Phoswich** & **VETO** detectors.

- **YAP(Ce)** performs good temp. dependence and **stability**.
- This system reduces B.G. down to **~ 1 %**.
- B.G. rate : **0.012** [counts / MeV / day]
~ 10⁻² [counts / MeV / day] (over 7.0 MeV)
- Not good PSD separation @ low energy (small pulse).
- High energy charged particle background induced by cosmic rays cannot be ignored.

Border for **Clear Evidence** : **3 counts / day F.G.**

Conclusion : Gas Permeation

Charged Particle Measurements were carried out
in **Gas Permeation Exp.** with **Pd Complexes**.

- **High Energy & Large Energy Loss particles**
were detected in **D₂ runs Only**.
99.97 % Reliability @ 7.0 – 9.5 MeV, Iwate sample
99.63 % Reliability @ 9.5 – 12.5 MeV, MHI sample
- Few counts caused by cosmic ray in **H₂ & Vac.** runs.

Detected particles are most likely to be ...

α Particles

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MHI

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